



# **Advanced Ionospheric Probe Onboard FORMOSAT-5 Satellite**

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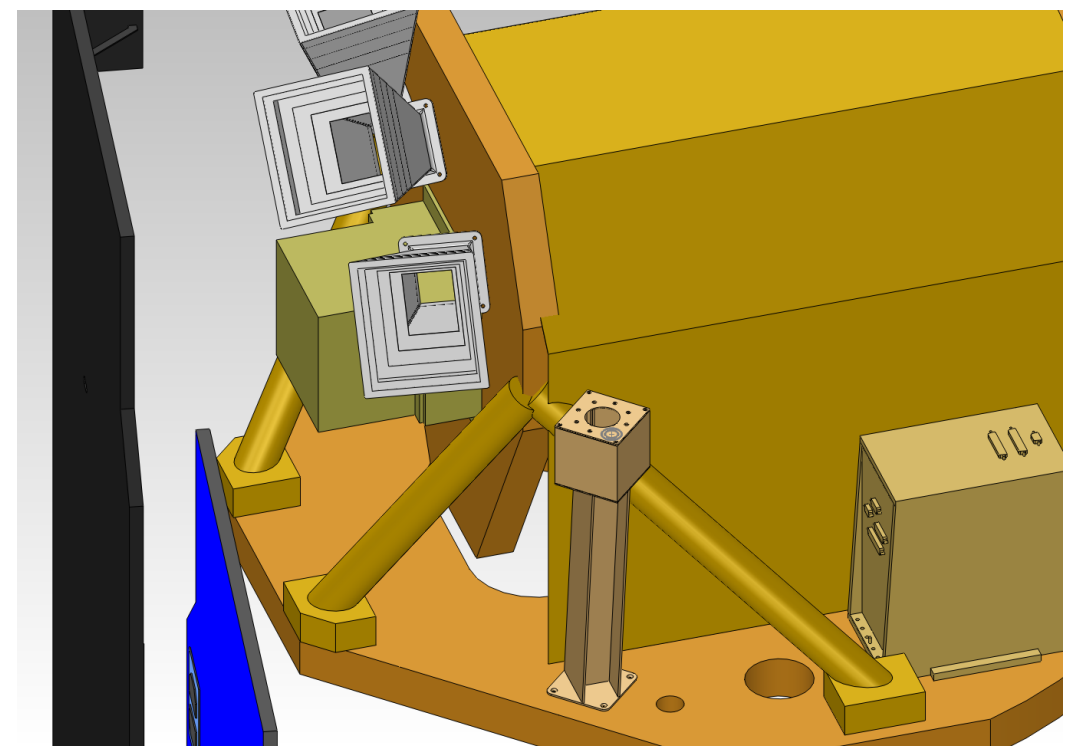
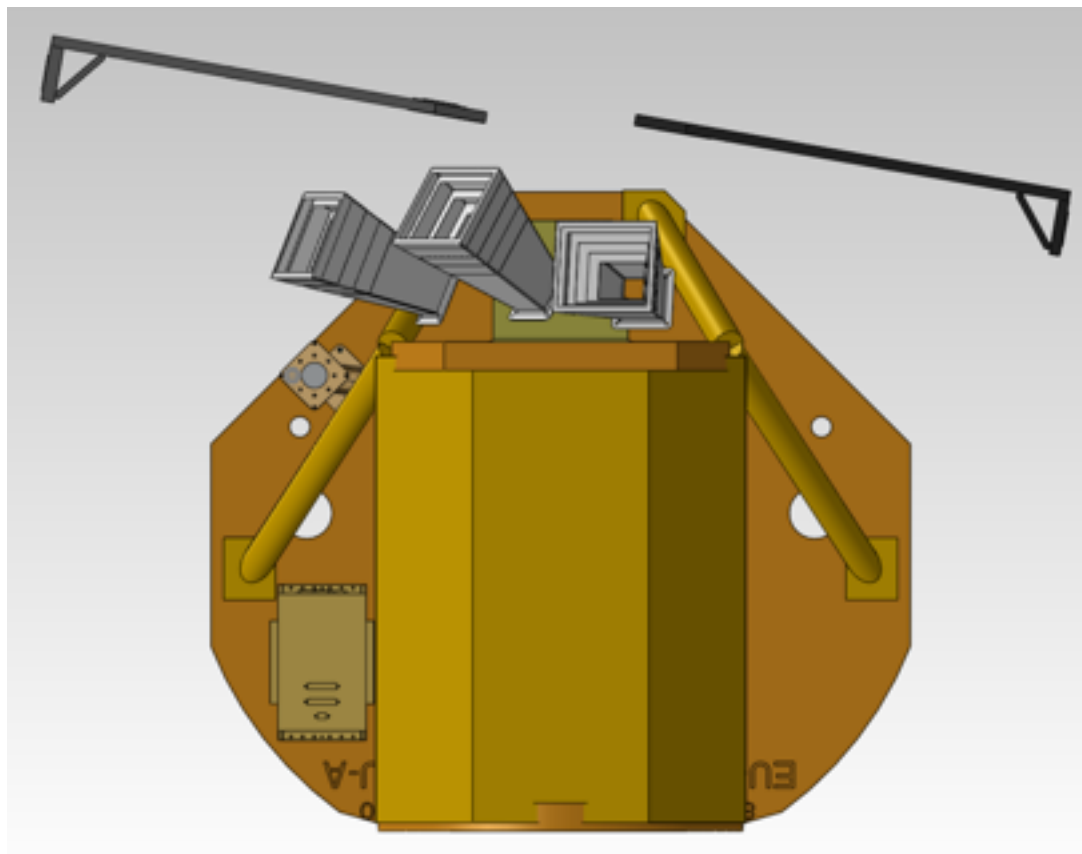
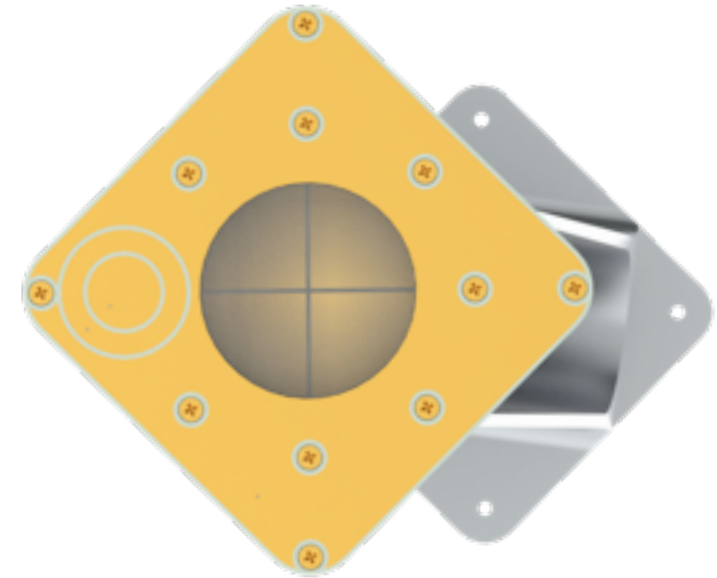
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UN/Japan Nano-satellite Symposium, Nagoya, Japan, 10-13 October, 2012

# Outline

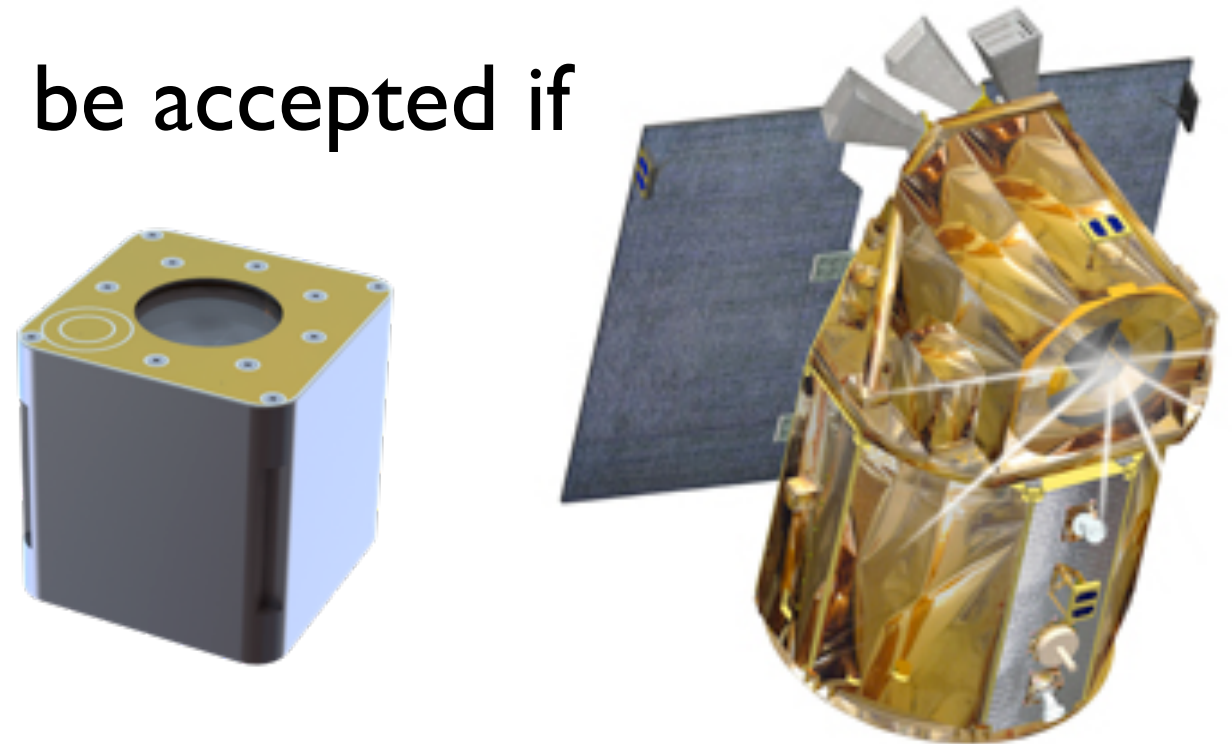
- Advanced Ionospheric Probe (AIP)
- Specifications
- Functional verifications
- Summary





# Advanced Ionospheric Probe

- FORMOSAT-5 mission orbit is anticipated at **720** km altitude with inclination **98.28°** as a sun-synchronous orbit in equatorial descending time between 9:45 am and 10:15 am LT.
- A piggyback science payload can be accepted if
  - Mass: < **5 kg**.
  - Power in average: < **5 W**.
  - Mission life time: > **2 years**.
- Advance Ionospheric Probe is an **all-in-one plasma sensor** to install on the FORMOSAT-5 satellite with sampling rate up to 8,192 Hz to measure ionospheric plasma concentrations ( **$N_i$** ), velocities ( **$V_i$** ), and temperatures ( **$T_i$**  and  **$T_e$** ) over a wide range of spatial scales in a time-sharing way.

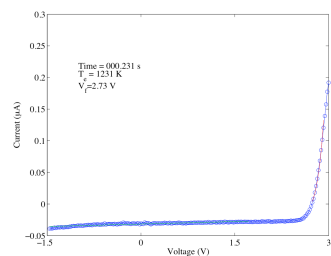




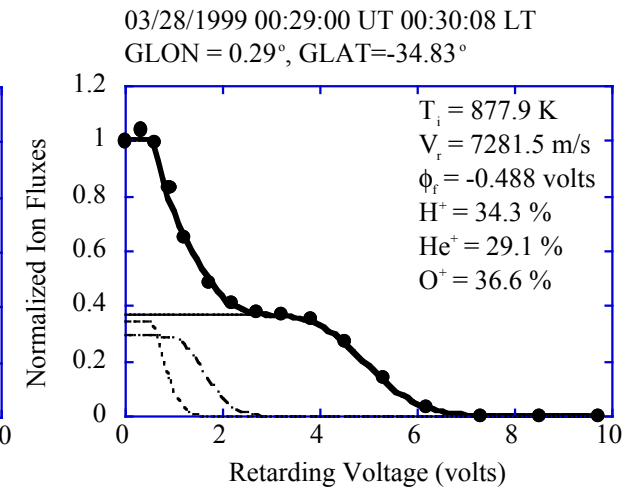
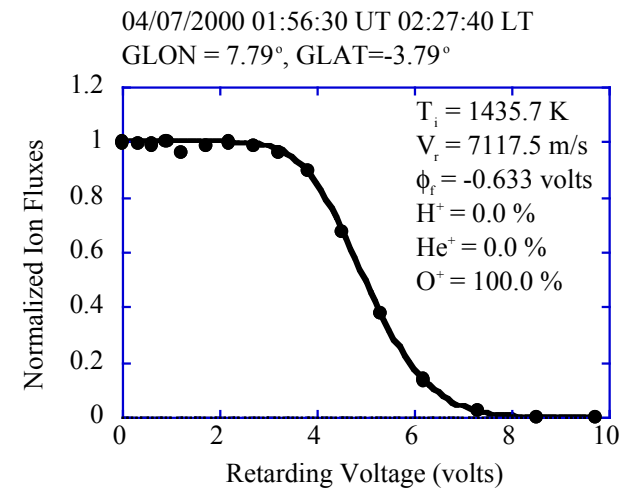
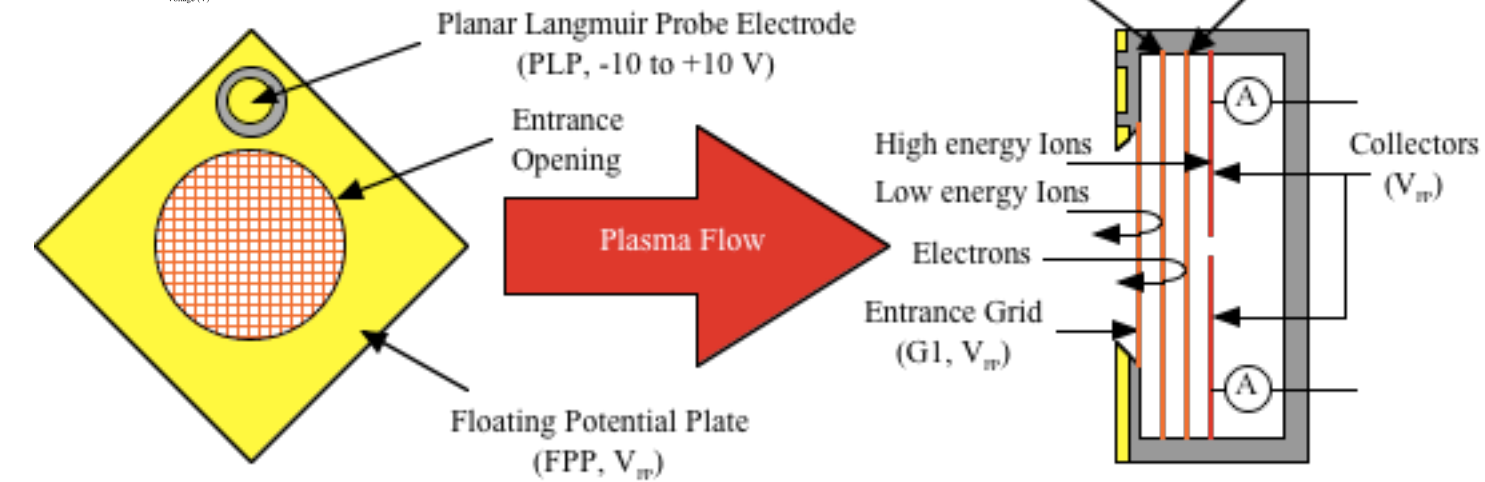


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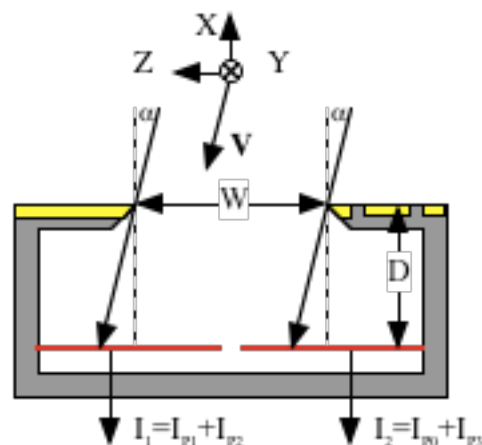
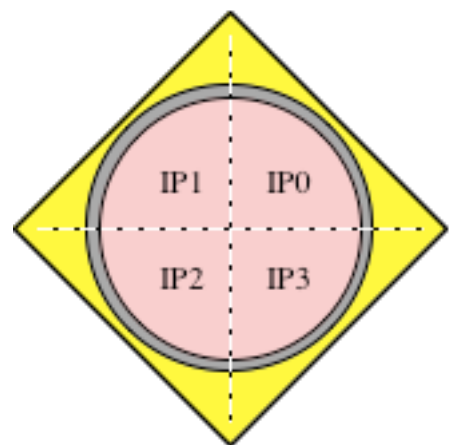
# Principle of measurement



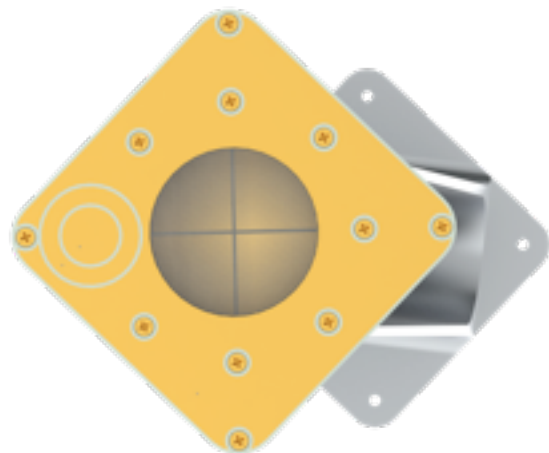
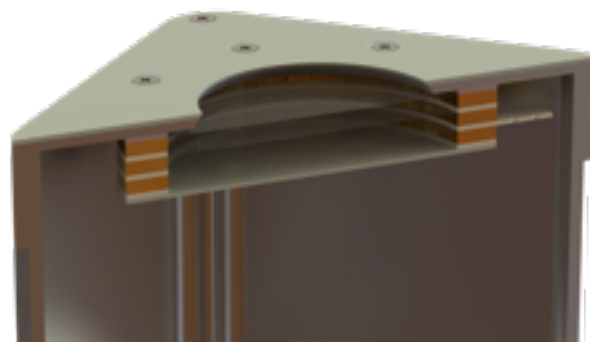
$$\frac{\kappa T_e}{e} = \frac{1}{\frac{d}{dV}[\ln(I)]}$$



$$I(V_M) = \sum_i I_i(V_M) = A \sum_i q_i n_{oi} U_i \left\{ \frac{1}{2} \left[ 1 + \operatorname{erf}(\beta_i F_i) + \frac{\exp[-(\beta_i F_i)^2]}{\sqrt{\pi} \beta_i U_i} \right] \right\}$$



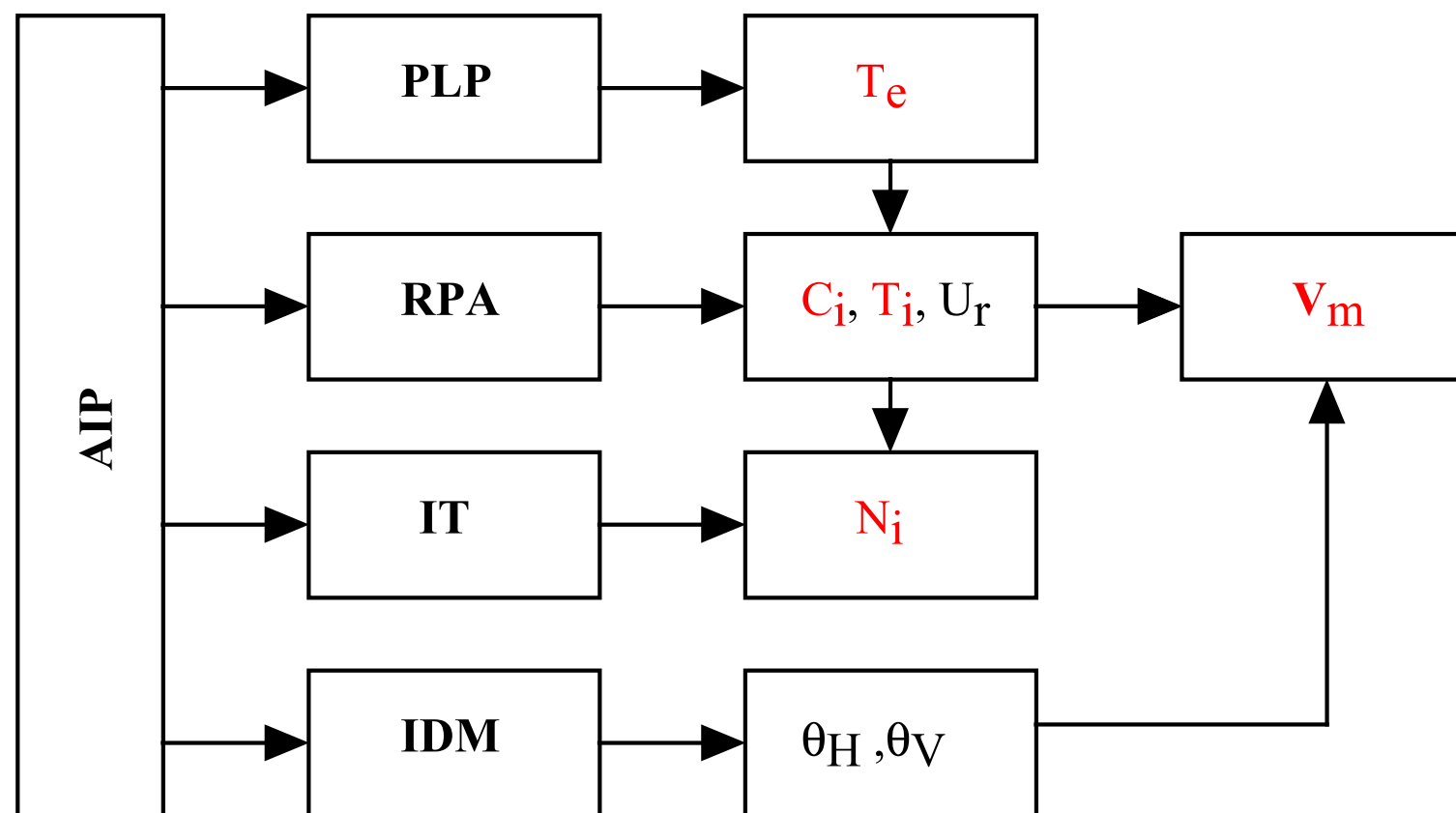
$$\tan \alpha = \left( \frac{W}{2D} \right) \left( \frac{I_1 - I_2}{I_1 + I_2} \right)$$



| Mode   | $N_i$    | Transverse components of the ion velocity | Others |
|--------|----------|-------------------------------------------|--------|
| NORMAL | 128 Hz   | 32 Hz                                     | 1 Hz   |
| FAST   | 1,024 Hz | 256 Hz                                    | 1 Hz   |
| BURST  | 8,192 Hz | 2,048 Hz                                  | 1 Hz   |



# Data flow and capabilities



| Parameters     | Range                                                          | Sensitivity                                             | Accuracy                                                |
|----------------|----------------------------------------------------------------|---------------------------------------------------------|---------------------------------------------------------|
| $C_i$          | 3% to 100%                                                     | 1%                                                      | 10%                                                     |
| $N_i$          | $4 \times 10^2$ to $1.2 \times 10^7 \text{ cm}^{-3}$           | 1%                                                      | 10%                                                     |
| $\mathbf{V}_i$ | $\pm 3.2 \text{ km s}^{-1}$<br>$\pm 5 \text{ km s}^{-1}$ (ram) | $\pm 10 \text{ m s}^{-1}$<br>$\pm 100 \text{ m s}^{-1}$ | $\pm 50 \text{ m s}^{-1}$<br>$\pm 200 \text{ m s}^{-1}$ |
| $T_i$          | 500 to 10,000 K                                                | $\pm 50 \text{ K}$                                      | $\pm 200 \text{ K}$                                     |
| $T_e$          | 500 to 10,000 K                                                | $\pm 50 \text{ K}$                                      | $\pm 200 \text{ K}$                                     |



FS-5/AIP

# Fine structures inside the equatorial spread F

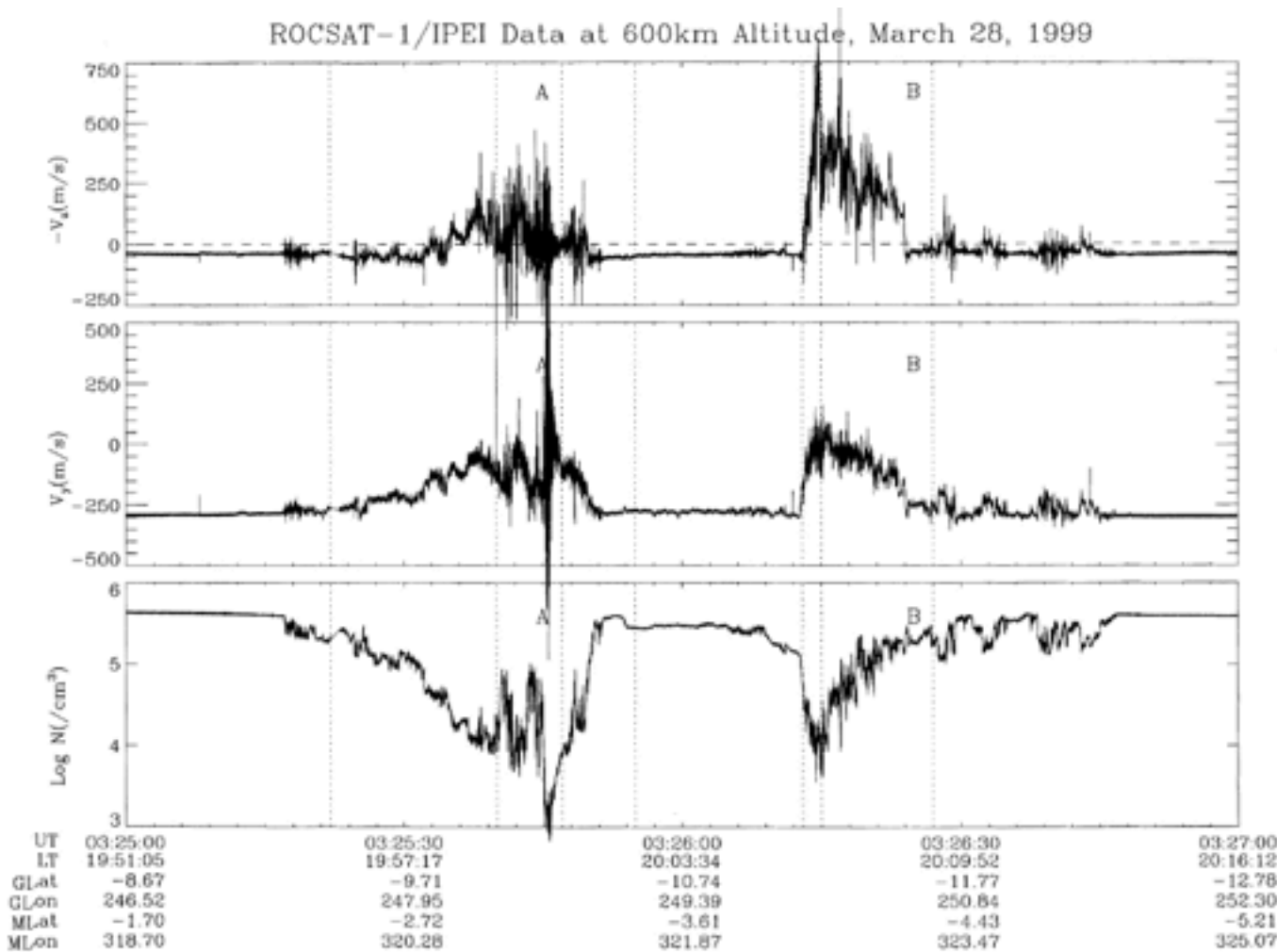


Figure 2. The expanded plot of Figure 1 for the time period 0325-0327 UT for bubble A and bubble B near the magnetic equator. The corotation component in  $V_z$  has been removed. Vertical dotted lines indicate the locations of sampled 1-s power spectra plotted in Figure 3.

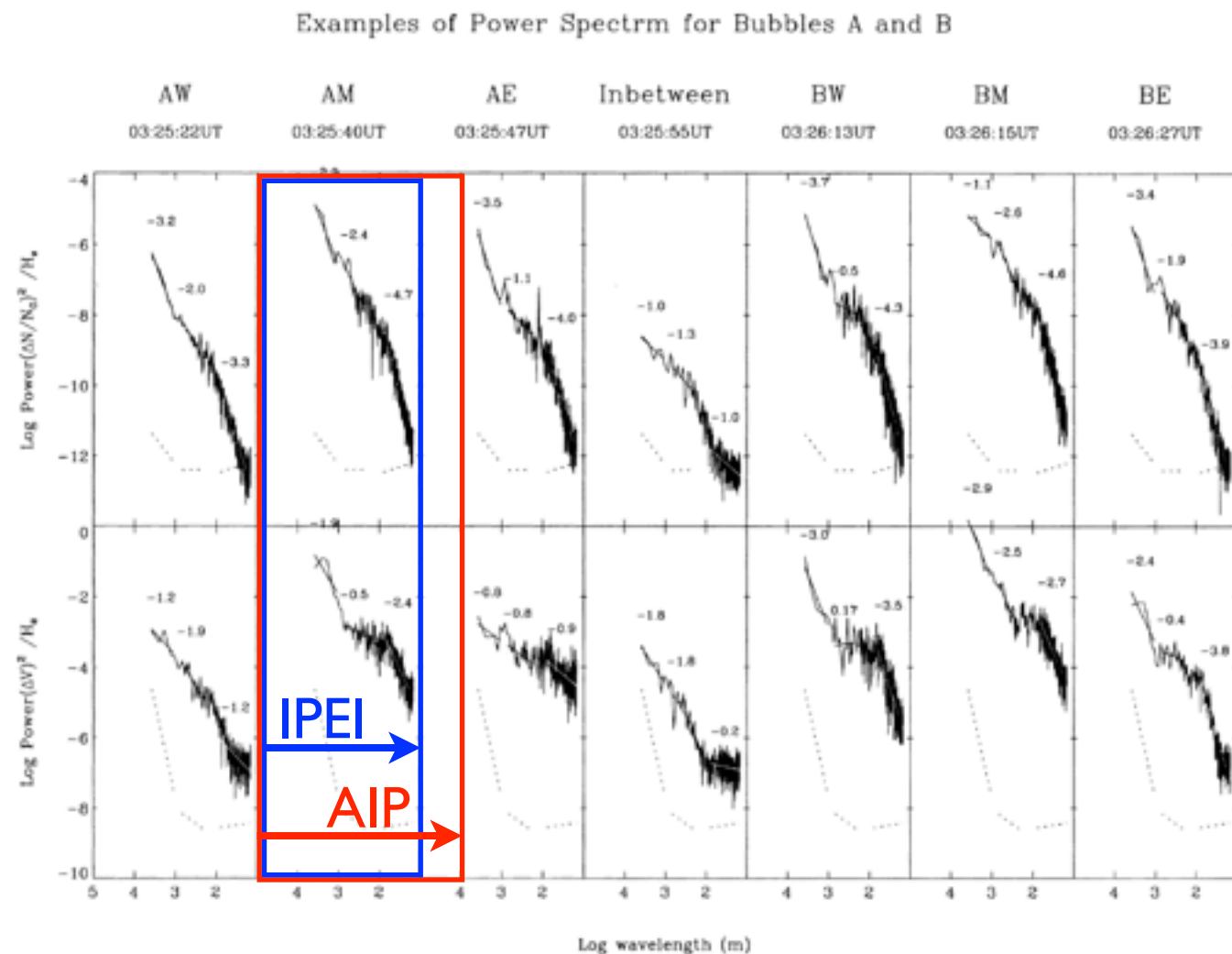


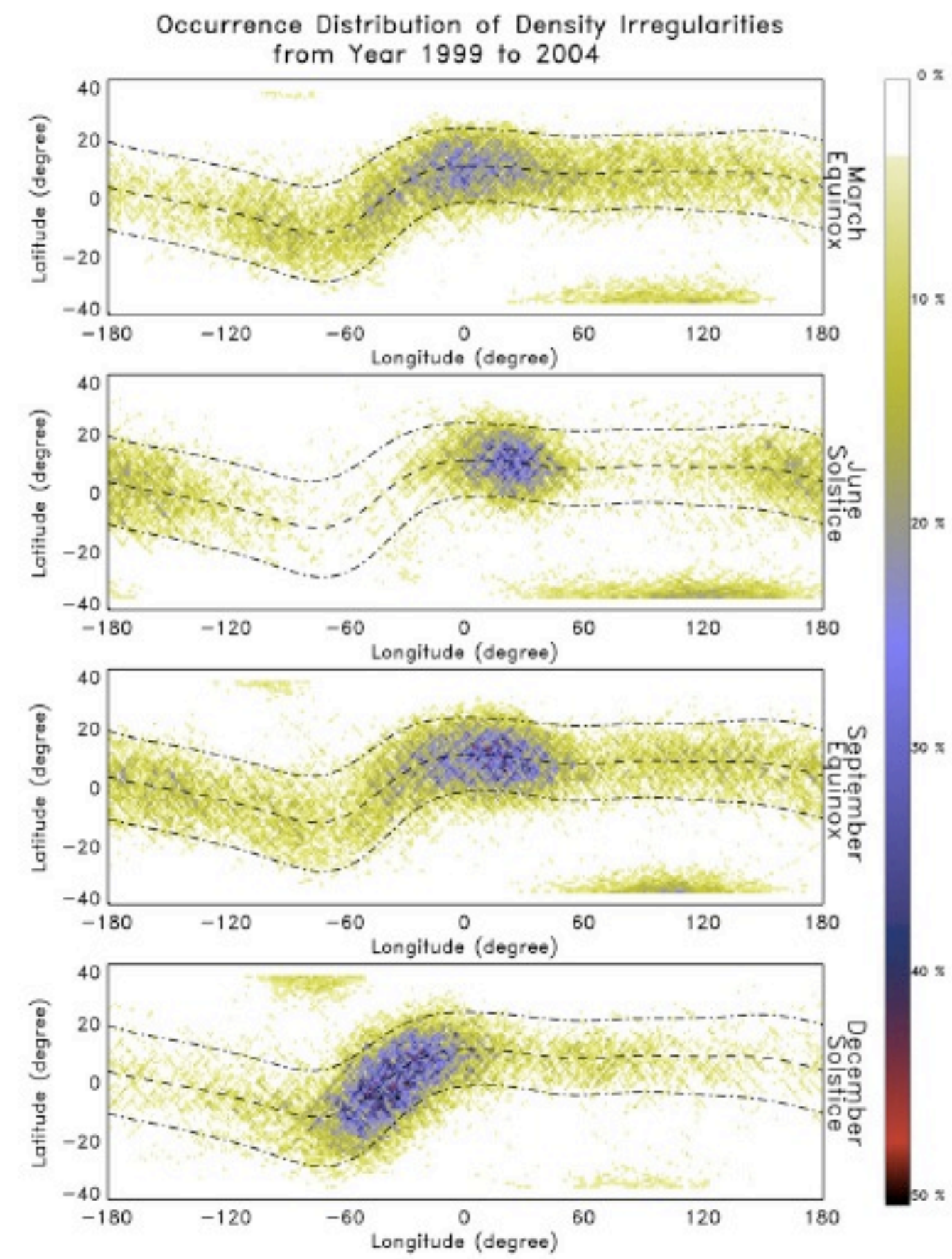
Figure 3. Examples of power spectra for  $(\Delta N/N_0)^2$  and  $\Delta V_z$  fluctuations at various parts of the bubbles as indicated by vertical lines in bubbles A and B in Figure 2.

Su et al., 2001, JGR

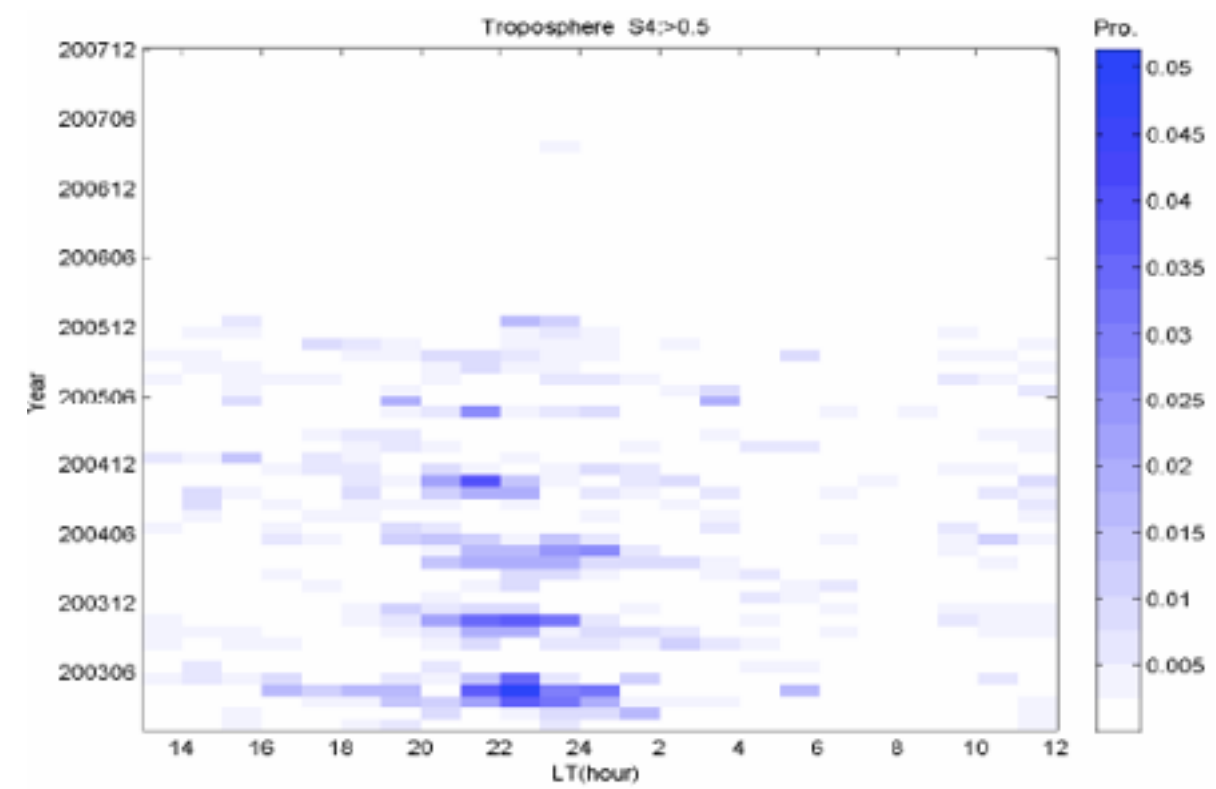
The AIP can explore 8 x finer scale fluctuations than those by ROCSAT-1/IPEI.



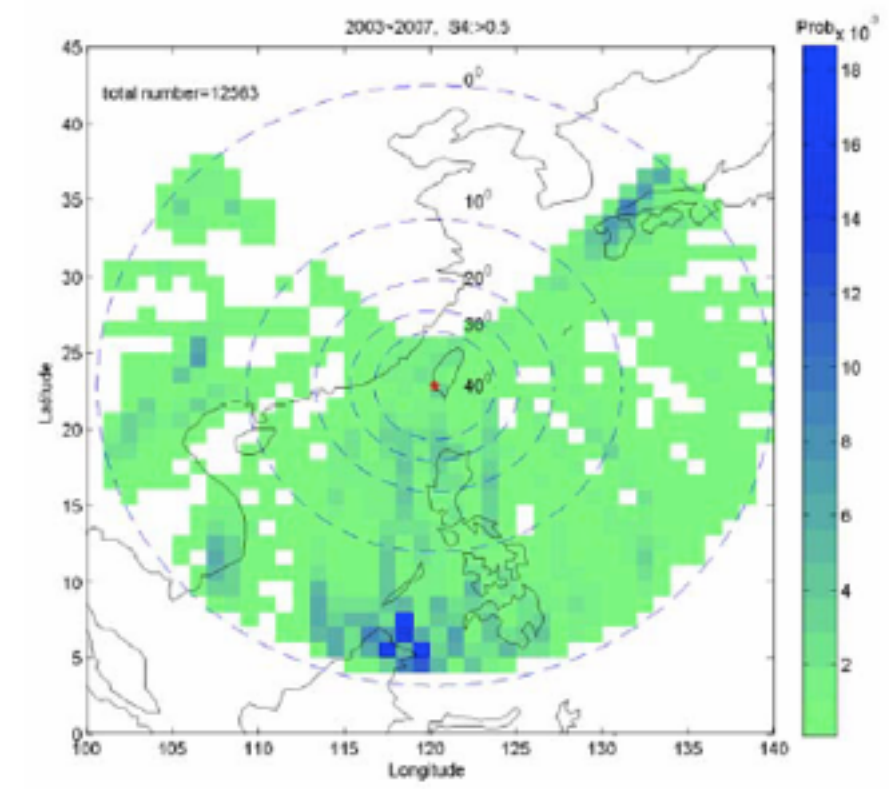
# Global observations on plasma irregularities



Su et al., 2008, JGR



Taiwan GPS scintillations ( $S_4 > 0.5$ ) caused by ESF

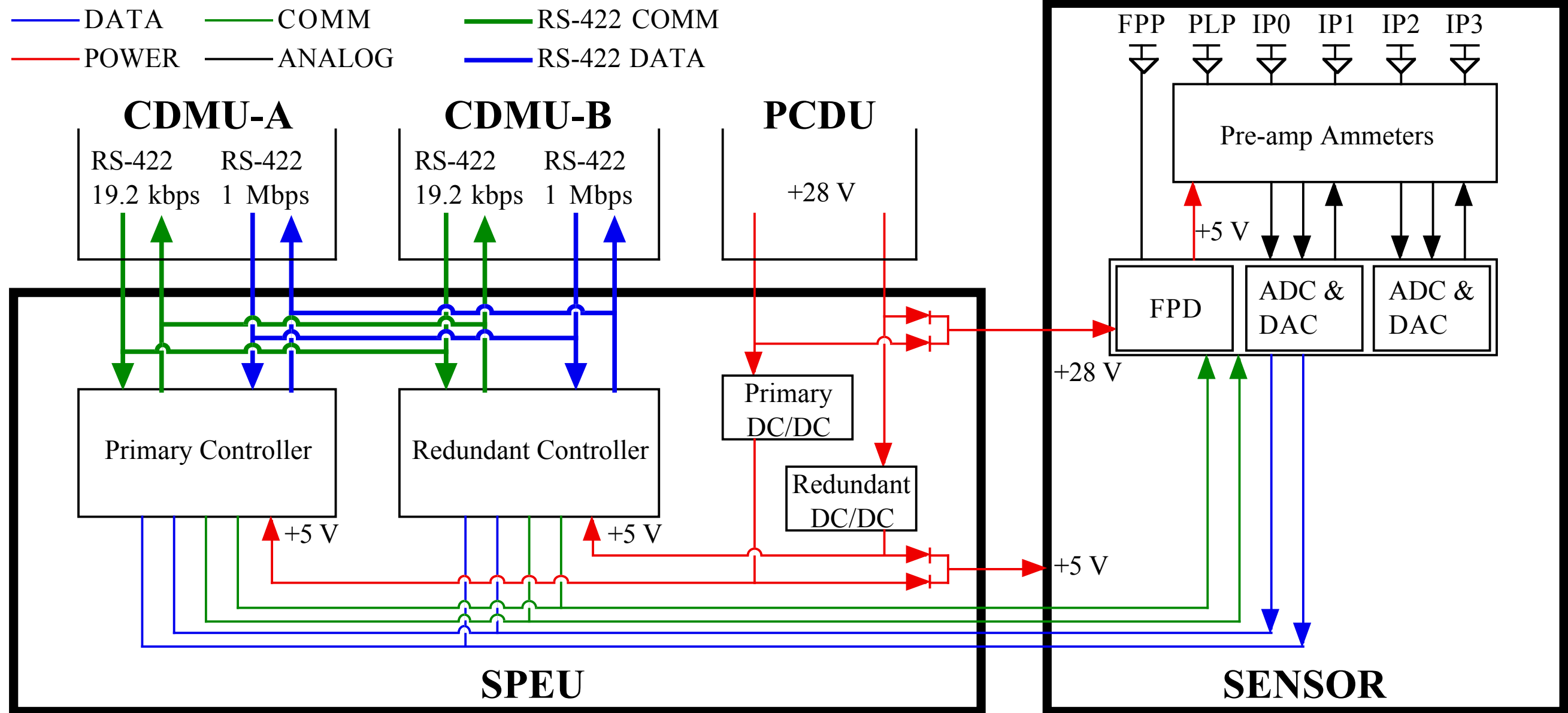


Chang, 2008, master thesis



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# Functional diagram



Redundant design for controllers, DC/DC, ADC and DAC, ammeters, etc.





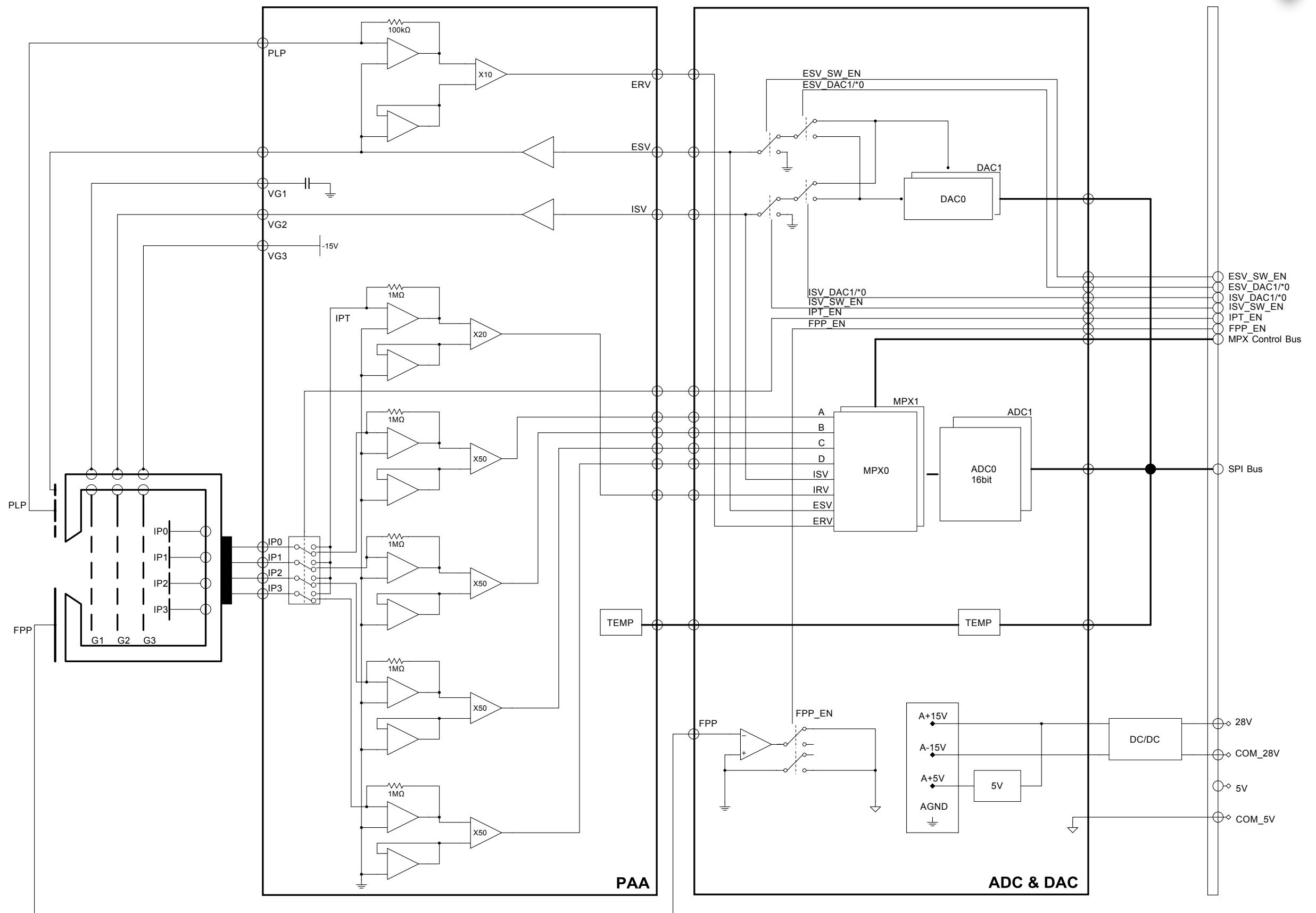
# Electrical specification

- Input voltage:  $28 \pm 6$  VDC ( $>$  the range between 26 V and 34 V).
- Power:  $\sim 4.74$  W in average
  - Sensor:  $2.24$  W (average).
  - SPEU:  $\sim 2.5$  W (average).
  - **Single event latch-up protection.**
- Dual command and telemetry interfaces (asynchronous RS-422, 19.2 kbps)
- Dual science data interfaces (synchronous RS-422, 1 Mbps)



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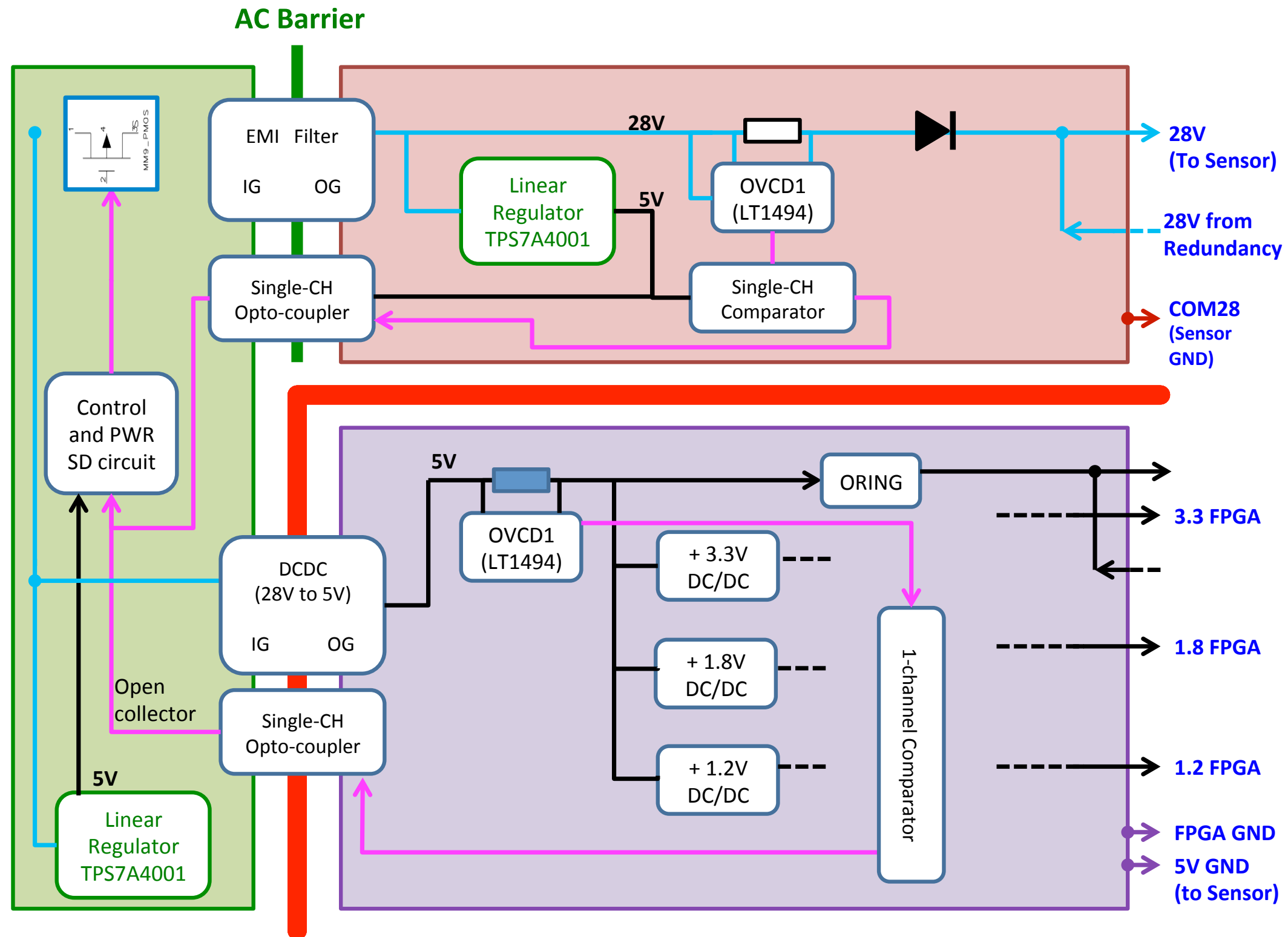
# Block diagram of the sensor





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# Single event latch-up protection

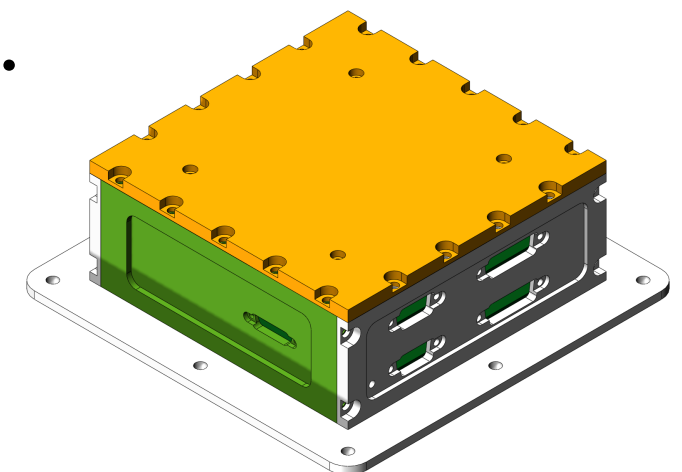
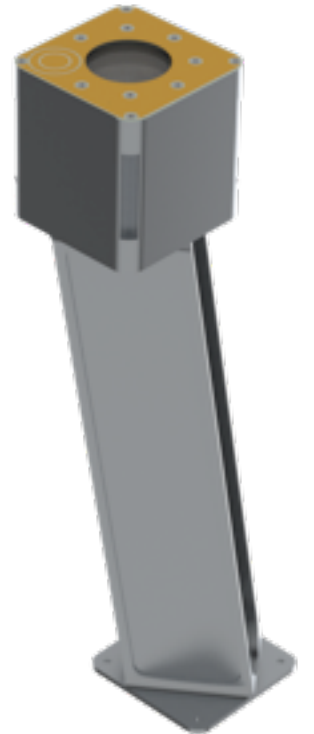






# Mechanical specification

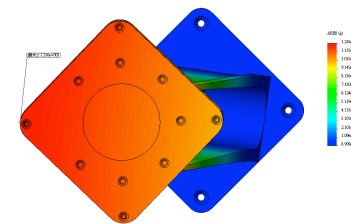
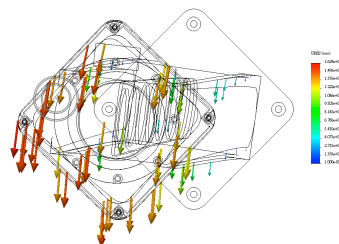
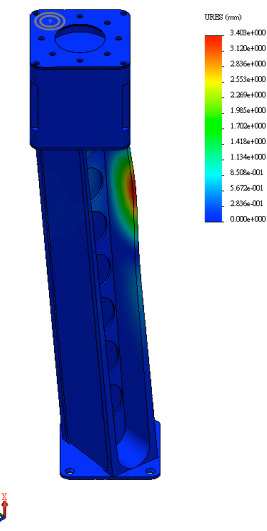
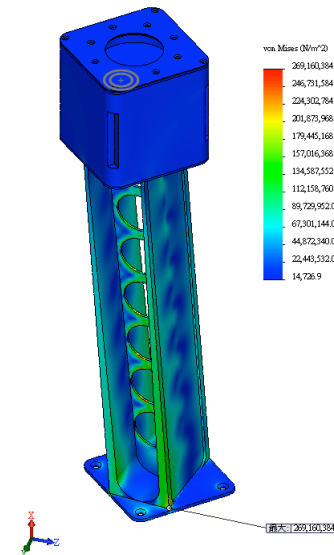
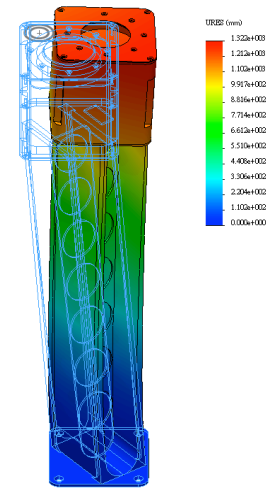
- Sensor: **1.46 kg** ( $< 1.5$  kg) and located on the top panel.
- Head:  $52^\circ$  field of view ( $> 50^\circ$ ).
  - Dimension: **100 (L) x 100 (W) x 100 mm (H)**.
  - Weight: **0.582 kg** (6061-T6).
- Stand:
  - Dimension: 140 (L) x 130 (W) x 350 mm (H).
  - Weight: 0.878 kg (7075-T6).
- SPEU: located inside S/C.
  - Dimension: 180 (L) x 180 (W) x 60 mm (H).
  - Weight:  $\sim 1.950$  kg (6061-T6).
- Harness from SPEU to sensor:  $< 1$  kg.
- Total mass:  $\sim$  **4.448 kg** ( $< 5$  kg).





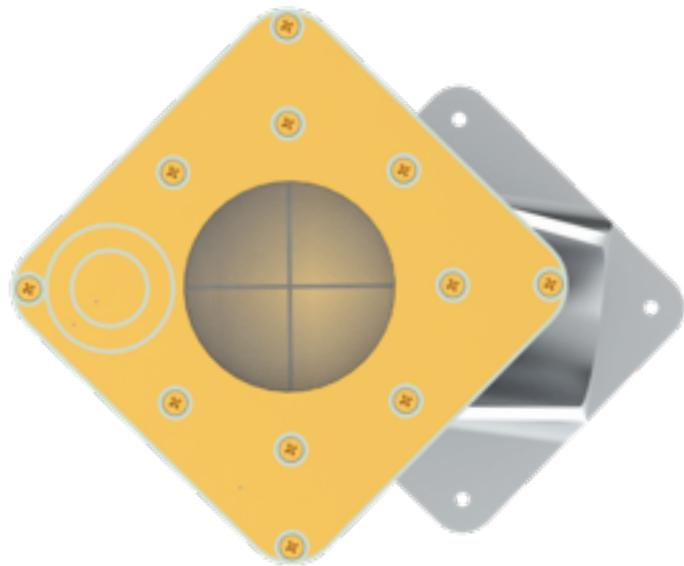
# Structure analysis for sensor

- Natural frequency: **145.6** Hz ( $> 120$  Hz).
- Load factors
  - Out of plane: **70** G ( $> 20$  G).
  - In-plane: **52** G ( $> 20$  G).
- Margin of safety (MS)
  - Ultimate: **2.10** ( $> 0$ ).
  - Yield: **2.72** ( $> 0$ ).
- Buckling safety margin ( $MS_B$ ): **3.45** ( $> 0$ ).
- Sinusoidal vibration (no collision during vibration)
  - 9G in-plane.
  - 10-300 Hz.



# Functional verifications

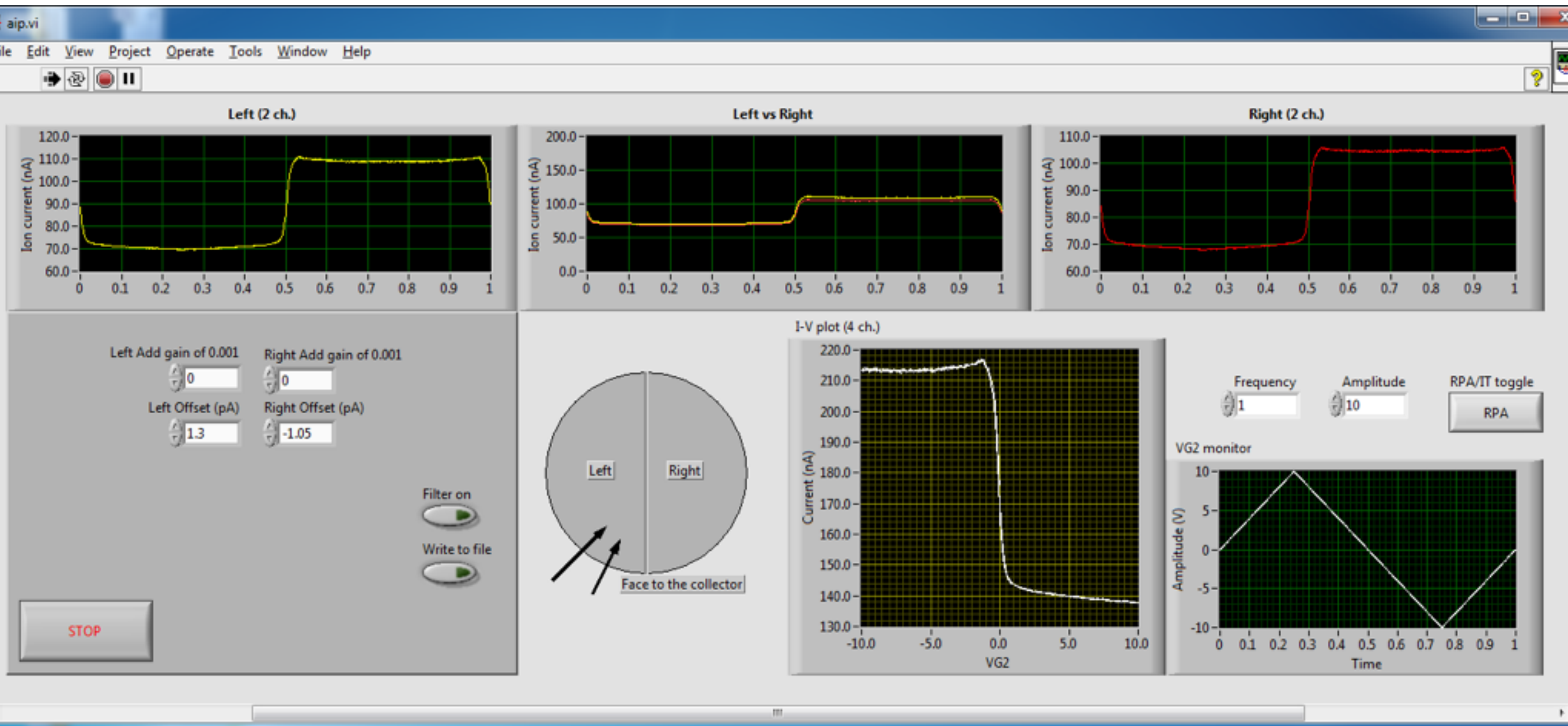
- IT submode
- RPA submode
- IDM submode
- ISD test





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# RPA mode (1 Hz)



AIP test on RPA mode with  $IRV = A+B+C+D$  (post-sum)

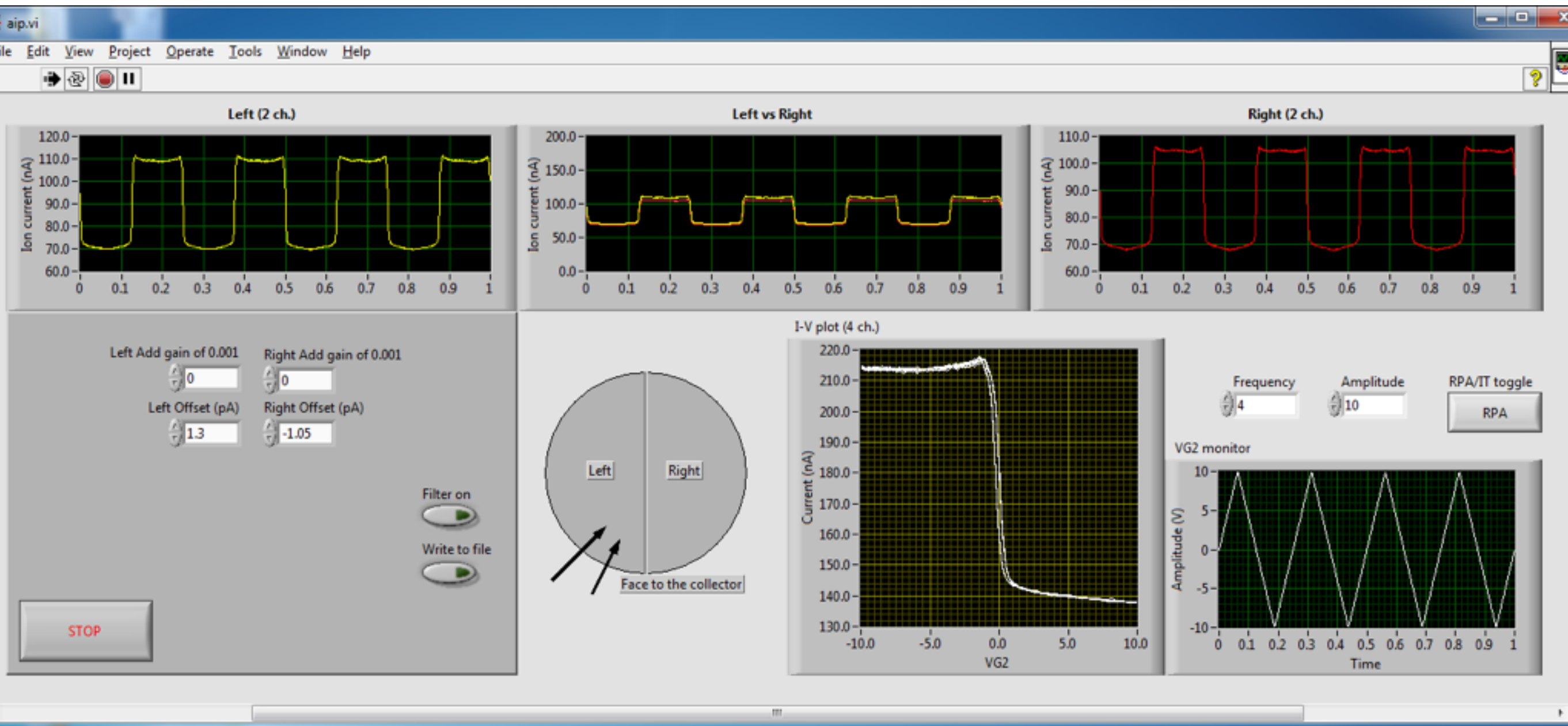
$$T_{i\_total} \sim T_{i\_A} \sim T_{i\_B} \sim T_{i\_C} \sim T_{i\_D} \sim 2,000 \text{ K}$$





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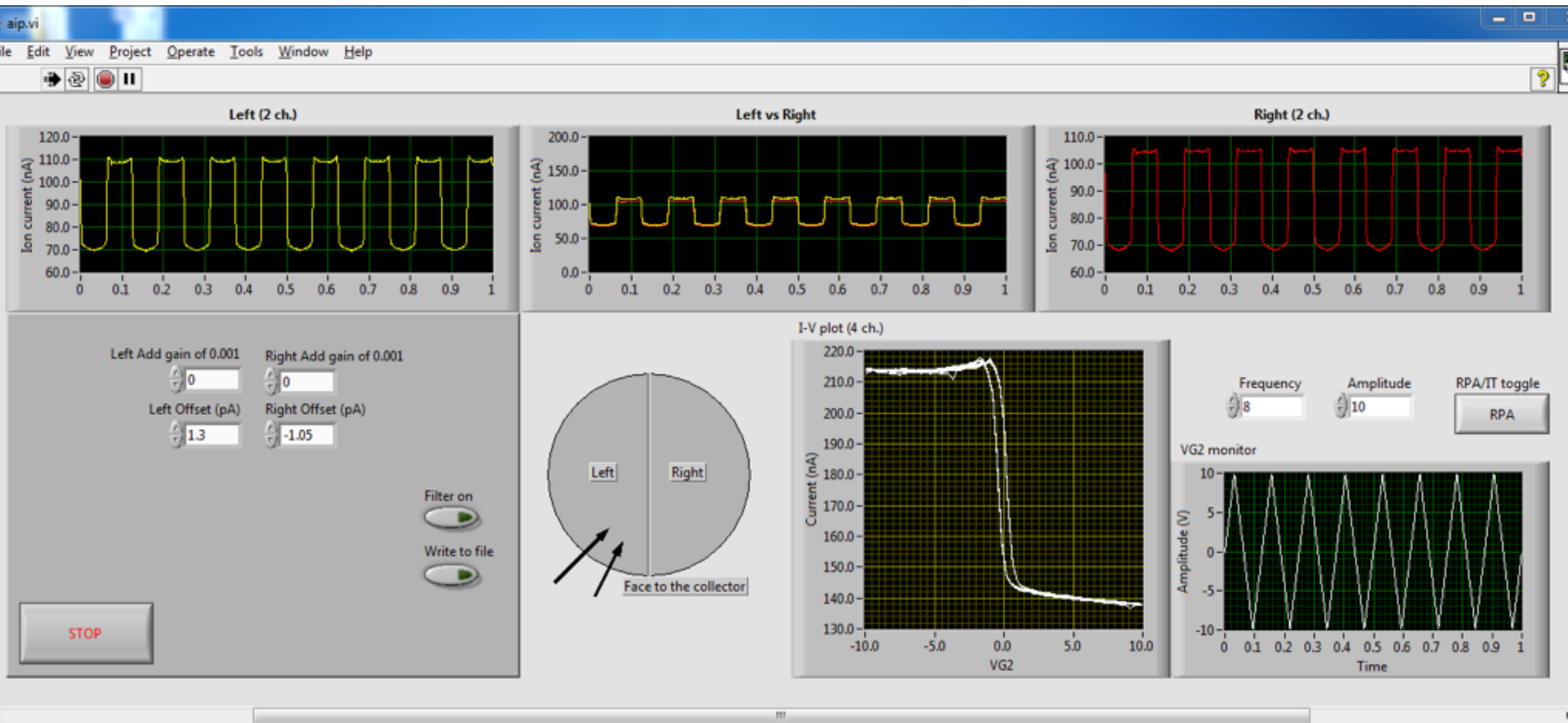
# RPA mode (4 Hz)





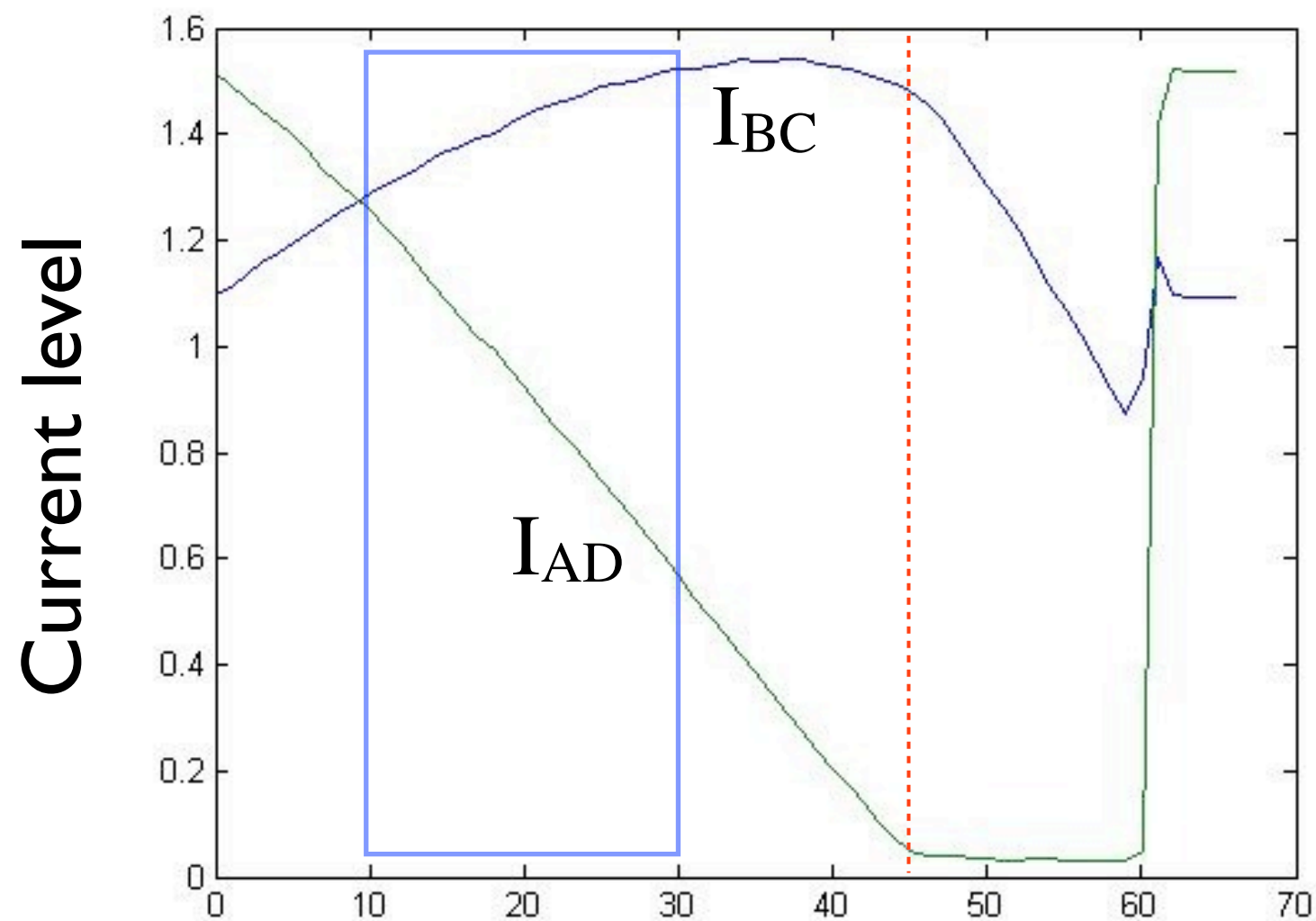
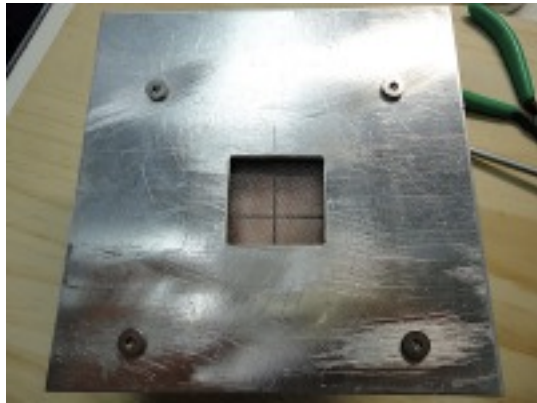
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# RPA mode (8 Hz)



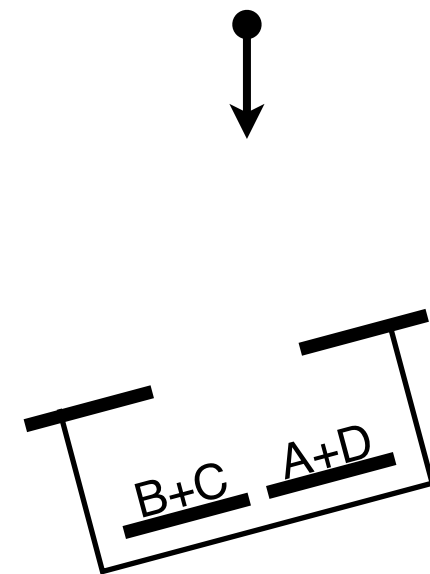
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# Ion drift meter mode

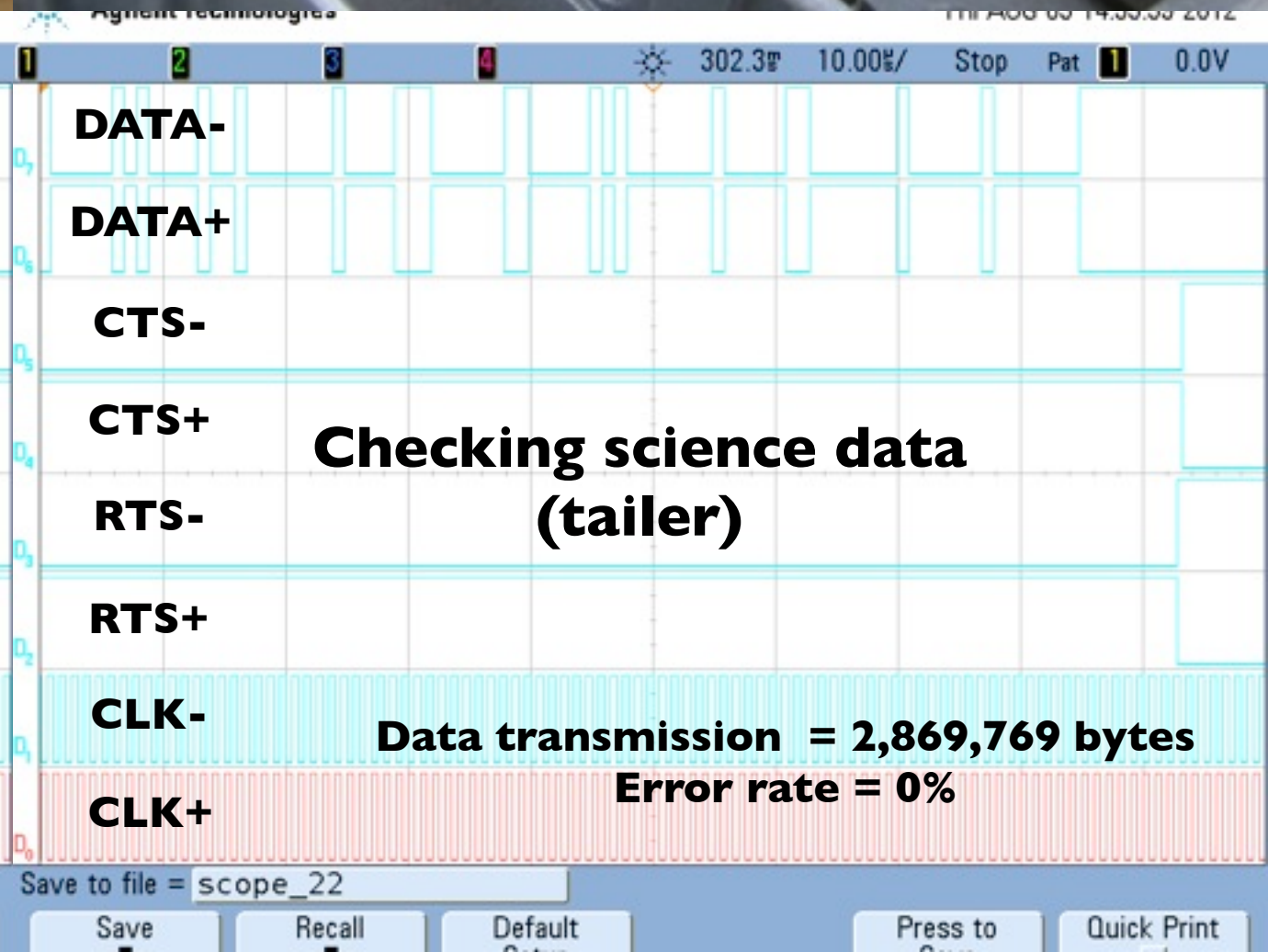
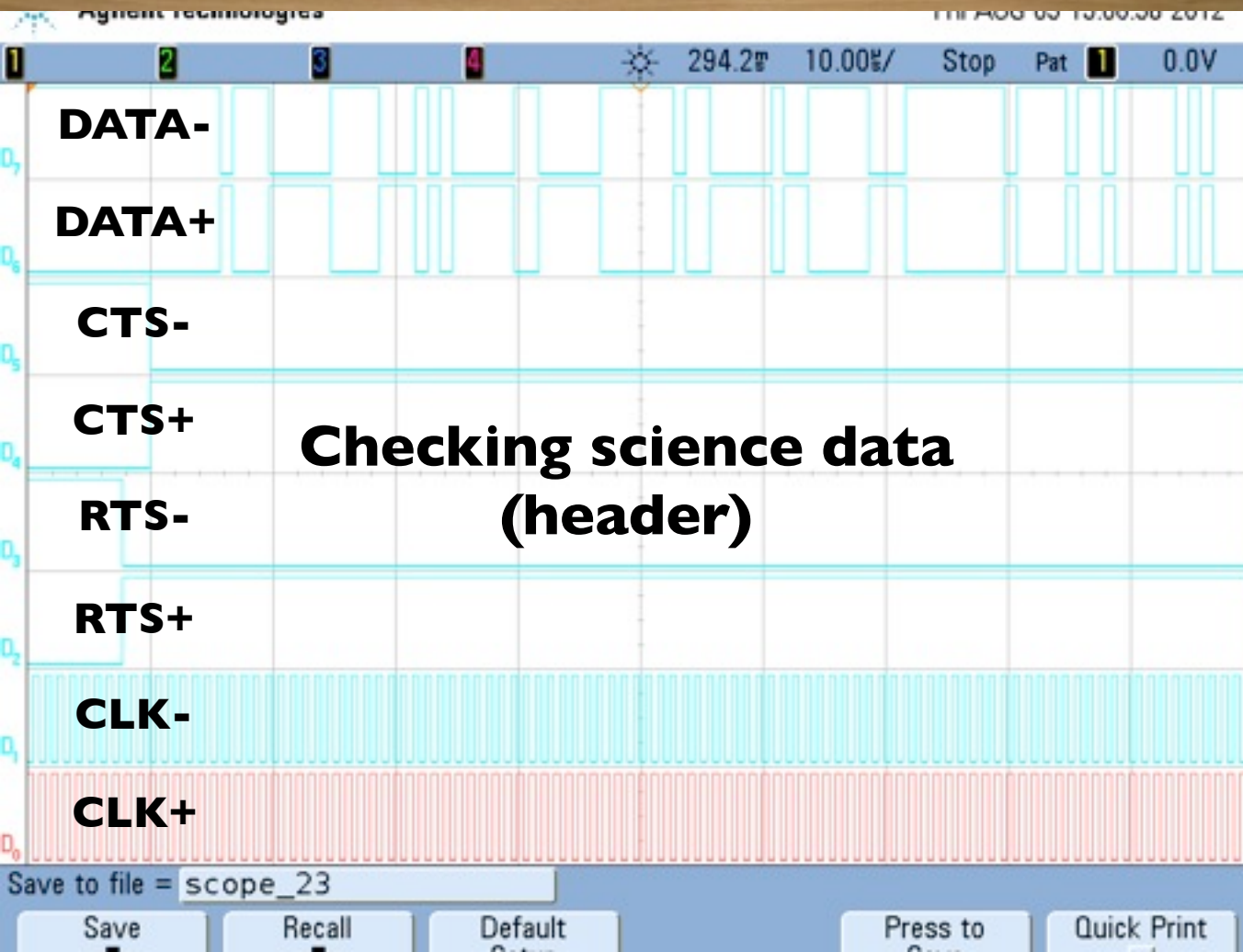
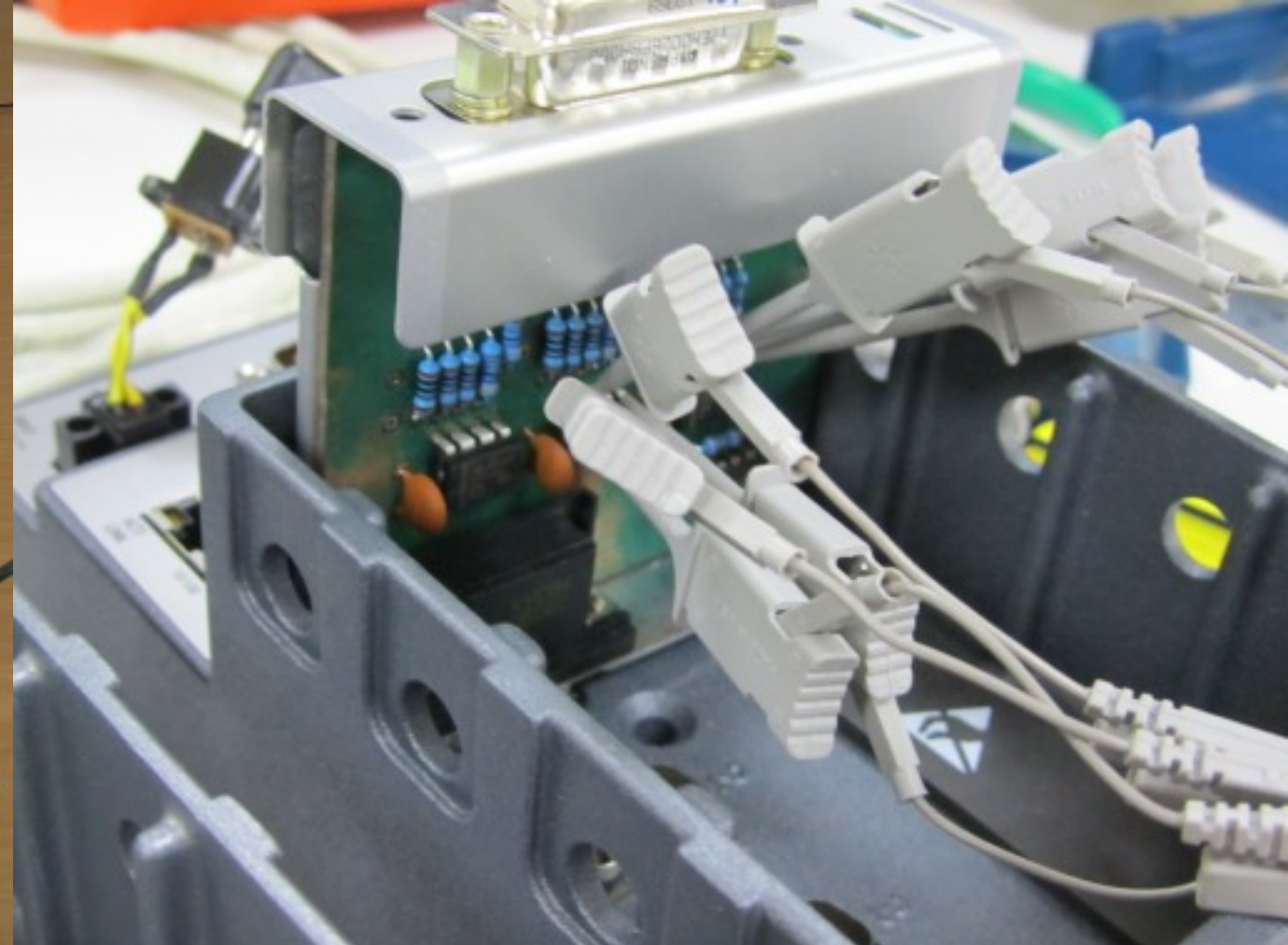


Time (second) or sensor angle (degree)

Plasma source











# Summary

- AIP progress: **WSD** (1/13/2012) → **MDR & SDR** (3/6/2012) → **PDR** (5/9/2012) → **ISD** (7/10/2012) → **CDR** (9/7/2012) → **DRR** (1/12/2013) → **PAR** (8/12/2013) ...
- Electrical interfaces and mechanical structure analysis are done.
- A sweeping frequency (1 Hz) is selected for all modes to reduce quasi-hysteresis effect in the I-V curves.
- In the future, the AIP can also be applied to complement GPS radio occultation experiment with in-situ measurements in a constellation (from 7th to 12th) of the FORMOSAT-7 micro-satellites. It is also possible to fit in a nano-satellite configuration to provide complete parameters of ionospheric plasma.